A Hybrid Model for Stellar Migration and Chemical Evolution in the Milky Way Disk

James W. Johnson 2021 SDSS Collaboration Meeting Tuesday, August 17, 2021







Context: Age-Metallicity, Age-Alpha Relations



Young, α -rich stars in the solar neighborhood

- "Rejuvenated" old stars? (Jofré++16, Yong++16, Izzard++18, Silva-Aguirre++18, Miglio++21)
- Not all of them (Hekker+Johnson19)

Age-Metallicity Relation non-monotonic (Feuillet++18)

• Migration of old, metal-rich stars born at small *R_{gal}* (Feuillet++18, Feuillet++19, Lu++21)



Left: Silva-Aguirre++18 Fig. 10; Right: Feuillet++18 Fig. 3

Multi-Zone + Hydro = Hybrid

Stellar Population forms at radius R, time T

- 1. Find *analog* that formed at similar R, T from *h*277 (Christensen++2012 suite, N-body Shop)
- 2. Assume analog's ΔR and present-day |z|
- 3. Migrate to final R with $\sqrt{\tau}$ -dependence
- 4. Eject nucleosynthetic products *as they migrate*

Assume SFH, $\dot{\Sigma}_{\star} - \Sigma_g$ relation, yields, outflows, then compute abundances

We use dR = 100 pc rings from R = 0 to 20 kpc ran for T = 13.2 Gyr, ~2.1 million stellar populations

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vice 1.2.1

pip install vice 🕒

After installing: python3 -m vice [--tutorial] [--docs]

Single Stellar Populations One-Zone Models Multi-Zone Models Calculate Yields

Versatile Integrator for Chemical Evolution

		NASA ADS	Griffith et al. (2021)
n Python 3.6 3.7 3.8 3.9 PyPI v	v1.2.1 License MIT		
G GitHub CI passing docs passing issue	ues 0 open		

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VICE is a user-friendly python package designed to model chemical enrichment in galaxies.

- 77 elements on the periodic table
- Fast integration of one-zone models
- Enrichment from single stellar populations
- Highly flexible nucleosynthetic yield calculations
- User-defined mathematical forms describing:
 - Nucleosynthetic yields in simulations
 - Mixing processes in multi-zone models
 - Infall and star formation histories
 - The stellar initial mass function
 - The star formation law
 - Element-by-element infall metallicities
 - Type Ia supernova delay-time distributions

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Age-[α /Fe]: The Impact of Stellar Migration

Model predicts young, α -enhanced stars in the solar neighborhood that migrated from large R



Under these models the SN Ia rate fluctuates, causing Fe-poor/rich ISM. Stars inherit these compositions, then migrate to inward



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The Impact of the SFH

The starburst models:

- 1. Predict substantial α -enhancement for young stars
- 2. Better reproduce C-shaped age-[O/H] relation

Different observables favor different models







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- Born out of an Fe-poor ISM caused by loss of SN Ia progenitors to migration
- No conflict with "rejuvenated" stars interpretation
 - Hekker+Johnson19 argue some of these stars are *intrinsically* young and α -rich



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- Enhances the C-shape of the age-metallicity relation
- Over-predicts $[\alpha/Fe]$ of young stars



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Additionally: (arxiv:2103.09838)

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VICE is publicly available! (pip install vice)

- If it can't handle your model, I'm interested in updating it so that it can
- Don't hesitate to ask for help

Join us on Slack!



The Origin of the Solar System Elements





Graphic created by Jennifer Johnson http://www.astronomy.ohio-state.edu/~jaj/nucleo/ Astronomical Image Credits: ESA/NASA/AASNova